The linear instability of some Einstein metrics

Changliang Wang

Max Planck Institute for Mathematics in Bonn

October 07 2019

Dirac operators in differential geometry and global analysis in memory of Professor Thomas Friedrich

Joint work with Uwe Semmelmann and McKenzie Wang



A Riemannian manifold (M^n, g) is Einstein if the Ricci curvature Ric_g is constant, i.e.

$$Ric_g = \Lambda g$$
 (1)

for some constant Λ , and Λ is called Einstein constant.

A Riemannian manifold (M^n, g) is Einstein if the Ricci curvature Ric_g is constant, i.e.

$$Ric_g = \Lambda g$$
 (1)

for some constant Λ , and Λ is called Einstein constant.

Let M^n be an n-dimensional compact manifold.

A Riemannian manifold (M^n, g) is Einstein if the Ricci curvature Ric_g is constant, i.e.

$$Ric_g = \Lambda g$$
 (1)

for some constant Λ , and Λ is called Einstein constant.

Let M^n be an n-dimensional compact manifold.

 $\mathcal{M} := \{ \text{Riemannian metrics } g \text{ on } M^n \}.$

A Riemannian manifold (M^n, g) is Einstein if the Ricci curvature Ric_g is constant, i.e.

$$Ric_g = \Lambda g$$
 (1)

for some constant Λ , and Λ is called Einstein constant.

Let M^n be an n-dimensional compact manifold.

 $\mathcal{M} := \{ \text{Riemannian metrics } g \text{ on } M^n \}.$

The normalized total scalar curvature functional is defined as

$$\widetilde{\mathbf{S}}: \mathcal{M} \to \mathbb{R}, \quad \widetilde{\mathbf{S}}(g) = \frac{1}{V(g)^{\frac{n-2}{n}}} \int_{\mathcal{M}} s_g d\mathrm{vol}_g,$$
 (2)

where s_g denotes the scalar curvature of g.

A Riemannian manifold (M^n, g) is Einstein if the Ricci curvature Ric_g is constant, i.e.

$$Ric_g = \Lambda g$$
 (1)

for some constant Λ , and Λ is called Einstein constant.

Let M^n be an n-dimensional compact manifold.

 $\mathcal{M} := \{ \text{Riemannian metrics } g \text{ on } M^n \}.$

The normalized total scalar curvature functional is defined as

$$\widetilde{\mathbf{S}}: \mathcal{M} \to \mathbb{R}, \quad \widetilde{\mathbf{S}}(g) = \frac{1}{V(g)^{\frac{n-2}{n}}} \int_{\mathcal{M}} s_g d\mathrm{vol}_g,$$
 (2)

where s_g denotes the scalar curvature of g.

The functional $\widetilde{\mathbf{S}}$ is diffeomorphism invariant and scaling invariant.



A Riemannian manifold (M^n, g) is Einstein if the Ricci curvature Ric_g is constant, i.e.

$$Ric_g = \Lambda g$$
 (1)

for some constant Λ , and Λ is called Einstein constant.

Let M^n be an n-dimensional compact manifold.

 $\mathcal{M} := \{ \text{Riemannian metrics } g \text{ on } M^n \}.$

The normalized total scalar curvature functional is defined as

$$\widetilde{\mathbf{S}}: \mathcal{M} \to \mathbb{R}, \quad \widetilde{\mathbf{S}}(g) = \frac{1}{V(g)^{\frac{n-2}{n}}} \int_{\mathcal{M}} s_g d\mathrm{vol}_g,$$
 (2)

where s_{ε} denotes the scalar curvature of g.

The functional $\tilde{\mathbf{S}}$ is diffeomorphism invariant and scaling invariant.

From now on, assume $n \ge 3$.



The first variation formula

Let g(t) for $t \in (-\tau, \tau)$ be a smooth family of metrics on M^n , g(0) = g, and $\frac{d}{dt}g(t)|_{t=0} = h \in C^{\infty}(M, T^*M \odot T^*M)$.

The first variation formula

Let g(t) for $t \in (-\tau, \tau)$ be a smooth family of metrics on M^n , g(0) = g, and $\frac{d}{dt}g(t)|_{t=0} = h \in C^{\infty}(M, T^*M \odot T^*M)$.

$$\widetilde{\mathbf{S}}_g' \cdot h := \frac{d}{dt} \widetilde{\mathbf{S}}(g(t))|_{t=0} = \frac{1}{V(g)^{\frac{n-2}{n}}} \int_M \langle -Ric_g + (\frac{s_g}{2} + \frac{2-n}{2n} \overline{s_g}) g, h \rangle d \mathrm{vol}_g, \tag{3}$$

where $\overline{s_g} = \frac{1}{V(g)} \int_M s_g dvol_g$ is the average of scaler curvature.

The first variation formula

Let g(t) for $t \in (-\tau, \tau)$ be a smooth family of metrics on M^n , g(0) = g, and $\frac{d}{dt}g(t)|_{t=0} = h \in C^{\infty}(M, T^*M \odot T^*M)$.

$$\widetilde{\mathbf{S}}_g' \cdot h := \frac{d}{dt} \widetilde{\mathbf{S}}(g(t))|_{t=0} = \frac{1}{V(g)^{\frac{n-2}{n}}} \int_M \langle -Ric_g + (\frac{s_g}{2} + \frac{2-n}{2n} \overline{s_g})g, h \rangle d\mathrm{vol}_g, \tag{3}$$

where $\overline{s_g} = \frac{1}{V(g)} \int_M s_g dvol_g$ is the average of scaler curvature.

Theorefore,

$$g$$
 is a critical point of $\widetilde{\mathbf{S}}\Leftrightarrow Ric_g=(rac{s_g}{2}+rac{2-n}{2n}\overline{s_g})g$ $\Leftrightarrow s_g$ is constant and g is Einstein (The second Bianchi identity, and $n\geq 3$.)

The second variation formula

At a such Einstein metric g,

$$\widetilde{\mathbf{S}}_{g}^{"}(h,h) := \frac{d^{2}}{dt^{2}}\widetilde{\mathbf{S}}(g(t))|_{t=0} = \frac{1}{V(g)^{\frac{n-2}{n}}} \int_{M} \langle P_{g}h, h \rangle d\mathrm{vol}_{g}, \tag{4}$$

The second variation formula

At a such Einstein metric g,

$$\widetilde{\mathbf{S}}_{g}^{"}(h,h) := \frac{d^{2}}{dt^{2}}\widetilde{\mathbf{S}}(g(t))|_{t=0} = \frac{1}{V(g)^{\frac{n-2}{n}}} \int_{M} \langle P_{g}h, h \rangle d\mathrm{vol}_{g}, \tag{4}$$

where

$$P_{g}h = -\frac{1}{2}\nabla^{*}\nabla h + \mathring{R}h + \delta_{g}^{*}(\delta_{g}h) + \frac{1}{2}\operatorname{Hess}_{g}(\operatorname{tr}_{g}h)$$

$$+ \left[-\frac{1}{2}(\Delta_{g}(\operatorname{tr}_{g}h)) + \frac{1}{2}\delta_{g}(\delta_{g}h) - \frac{s_{g}}{2n}(\operatorname{tr}_{g}h) \right]g$$

$$- \frac{(2-n)s_{g}}{2n^{2}} \overline{(\operatorname{tr}_{g}h)}g,$$

$$(5)$$

 $\overline{(\operatorname{tr}_{g} h)} = \frac{1}{V(g)} \int_{M} (\operatorname{tr}_{g} h) d\operatorname{vol}_{g}$, and $(\mathring{R}h)_{ij} = R_{ikjl} h_{kl}$.

The second variation formula

At a such Einstein metric g,

$$\widetilde{\mathbf{S}}_{g}^{"}(h,h) := \frac{d^{2}}{dt^{2}}\widetilde{\mathbf{S}}(g(t))|_{t=0} = \frac{1}{V(g)^{\frac{n-2}{n}}} \int_{M} \langle P_{g}h, h \rangle d\mathrm{vol}_{g}, \tag{4}$$

where

$$P_{g}h = -\frac{1}{2}\nabla^{*}\nabla h + \mathring{R}h + \delta_{g}^{*}(\delta_{g}h) + \frac{1}{2}\operatorname{Hess}_{g}(\operatorname{tr}_{g}h)$$

$$+ \left[-\frac{1}{2}(\Delta_{g}(\operatorname{tr}_{g}h)) + \frac{1}{2}\delta_{g}(\delta_{g}h) - \frac{s_{g}}{2n}(\operatorname{tr}_{g}h)\right]g$$

$$-\frac{(2-n)s_{g}}{2n^{2}}\overline{(\operatorname{tr}_{g}h)}g,$$

$$(5)$$

 $\overline{(\operatorname{tr}_{g} h)} = \frac{1}{V(g)} \int_{M} (\operatorname{tr}_{g} h) d\operatorname{vol}_{g}$, and $(\mathring{R}h)_{ij} = R_{ikjl} h_{kl}$.



$$C^{\infty}(M, T^*M \odot T^*M) = \operatorname{Im} \delta_g^* \oplus C^{\infty}(M) \cdot g \oplus (\ker \operatorname{tr}_g \cap \ker \delta_g).$$
 (6)

Let (M^n, g) be a compact Einstein manifold other than standard spheres. Then

$$C^{\infty}(M, T^*M \odot T^*M) = \operatorname{Im} \delta_g^* \oplus C^{\infty}(M) \cdot g \oplus (\ker \operatorname{tr}_g \cap \ker \delta_g).$$
 (6)

ullet This decomposition is orthogonal with respect to $\langle P_{\rm g} h, \tilde{h} \rangle_{L^2}.$

$$C^{\infty}(M, T^*M \odot T^*M) = \operatorname{Im} \delta_g^* \oplus C^{\infty}(M) \cdot g \oplus (\ker \operatorname{tr}_g \cap \ker \delta_g).$$
 (6)

- ullet This decomposition is orthogonal with respect to $\langle P_{g}h, \tilde{h} \rangle_{L^{2}}$.
- $\mathrm{Im}\delta_g^*$ corresponds to diffeomorphisms. Thus, $\langle P_g h, \tilde{h} \rangle_{L^2} = 0$, for $h \in \mathrm{Im}\delta_g^*$ and any \tilde{h} .

$$C^{\infty}(M, T^*M \odot T^*M) = \operatorname{Im} \delta_g^* \oplus C^{\infty}(M) \cdot g \oplus (\ker \operatorname{tr}_g \cap \ker \delta_g).$$
 (6)

- ullet This decomposition is orthogonal with respect to $\langle P_{\rm g} h, \tilde{h} \rangle_{L^2}.$
- $\mathrm{Im}\delta_g^*$ corresponds to diffeomorphisms. Thus, $\langle P_g h, \tilde{h} \rangle_{L^2} = 0$, for $h \in \mathrm{Im}\delta_g^*$ and any \tilde{h} .
- Moreover, $\langle P_g(fg), fg \rangle_{L^2} = \frac{n-2}{2} \int_M (-(n-1)\Delta_g f s_g f) \cdot f dvol_g \ge 0$, for any $f \in C^{\infty}(M)$.

$$C^{\infty}(M, T^*M \odot T^*M) = \operatorname{Im} \delta_g^* \oplus C^{\infty}(M) \cdot g \oplus (\ker \operatorname{tr}_g \cap \ker \delta_g).$$
 (6)

- This decomposition is orthogonal with respect to $\langle P_g h, \hat{h} \rangle_{L^2}$.
- $\mathrm{Im}\delta_g^*$ corresponds to diffeomorphisms. Thus, $\langle P_g h, \tilde{h} \rangle_{L^2} = 0$, for $h \in \mathrm{Im}\delta_g^*$ and any \tilde{h} .
- Moreover, $\langle P_g(fg), fg \rangle_{L^2} = \frac{n-2}{2} \int_M (-(n-1)\Delta_g f s_g f) \cdot f dvol_g \ge 0$, for any $f \in C^{\infty}(M)$.
- An Einstein metric is always a saddle point of the normalized total scalar curvature functional.

Definition of linear stability of Einstein metrics

In the direction orthogonal to diffeomorphism and conformal changes, i.e. $h \in \ker \operatorname{tr}_g \cap \ker \delta_g$,

$$\widetilde{\mathbf{S}}_{g}^{"}(h,h) = \frac{-1}{2V(g)^{\frac{n-2}{n}}} \int_{M} \langle \nabla^{*} \nabla h - 2\mathring{R}h, h \rangle d \mathrm{vol}_{g}. \tag{7}$$

Definition of linear stability of Einstein metrics

In the direction orthogonal to diffeomorphism and conformal changes, i.e. $h \in \ker \operatorname{tr}_g \cap \ker \delta_g$,

$$\widetilde{\mathbf{S}}_{g}^{"}(h,h) = \frac{-1}{2V(g)^{\frac{n-2}{n}}} \int_{M} \langle \nabla^{*} \nabla h - 2\mathring{R}h, h \rangle d \mathrm{vol}_{g}. \tag{7}$$

Einstein operator: $\nabla^*\nabla - 2\mathring{R} = \Delta_L - 2\Lambda$.

Definition of linear stability of Einstein metrics

In the direction orthogonal to diffeomorphism and conformal changes, i.e. $h \in \ker \operatorname{tr}_g \cap \ker \delta_g$,

$$\widetilde{\mathbf{S}}_{g}^{"}(h,h) = \frac{-1}{2V(g)^{\frac{n-2}{n}}} \int_{M} \langle \nabla^{*} \nabla h - 2\mathring{R}h, h \rangle d \mathrm{vol}_{g}. \tag{7}$$

Einstein operator: $\nabla^* \nabla - 2 \mathring{R} = \Delta_L - 2 \Lambda$.

Definition: A (compact) Einstein metric (M^n, g) is \widetilde{S} -linearly stable, if

$$\langle \nabla^* \nabla h - 2 \mathring{R} h, h \rangle_{L^2} \ge 0 \tag{8}$$

for all TT-tensors h, i.e. symmetric 2-tensor h satisfying $\operatorname{tr}_g h = 0$ and $\delta_g h = 0$, and otherwise, $\widetilde{\mathbf{S}}$ -linearly unstable.

Examples of \widetilde{S} -linearly stable Einstein metrics

• Einstein manifolds with negative sectional curvature (N. Koiso, 1978).

Examples of \widetilde{S} -linearly stable Einstein metrics

- Einstein manifolds with negative sectional curvature (N. Koiso, 1978).
- Many compact irreducible symmetric spaces, e.g. S^n , $\mathbb{C}P^n$, $\mathbb{H}P^n$, and all compact simple Lie groups except $\mathrm{Sp}(n)$, $n \geq 2$ and $\mathrm{SU}(n)$, $n \geq 3$ (N. Koiso, 1980).

- Einstein manifolds with negative sectional curvature (N. Koiso, 1978).
- Many compact irreducible symmetric spaces, e.g. S^n , $\mathbb{C}P^n$, $\mathbb{H}P^n$, and all compact simple Lie groups except $\mathrm{Sp}(n)$, $n \geq 2$ and $\mathrm{SU}(n)$, $n \geq 3$ (N. Koiso, 1980).
- Compact manifolds with non-zero parallel spinors (X. Dai, X. Wang, and G. Wei, 2005).

- Einstein manifolds with negative sectional curvature (N. Koiso, 1978).
- Many compact irreducible symmetric spaces, e.g. S^n , $\mathbb{C}P^n$, $\mathbb{H}P^n$, and all compact simple Lie groups except $\mathrm{Sp}(n)$, $n \geq 2$ and $\mathrm{SU}(n)$, $n \geq 3$ (N. Koiso, 1980).
- Compact manifolds with non-zero parallel spinors (X. Dai, X. Wang, and G. Wei, 2005). Therefore, so far all known examples of compact Ricci flat manifold are stable.

- Einstein manifolds with negative sectional curvature (N. Koiso, 1978).
- Many compact irreducible symmetric spaces, e.g. S^n , $\mathbb{C}P^n$, $\mathbb{H}P^n$, and all compact simple Lie groups except $\mathrm{Sp}(n)$, $n \geq 2$ and $\mathrm{SU}(n)$, $n \geq 3$ (N. Koiso, 1980).
- Compact manifolds with non-zero parallel spinors (X. Dai, X. Wang, and G. Wei, 2005). Therefore, so far all known examples of compact Ricci flat manifold are stable.
- Compact Einstein manifolds with non-positive scalar curvature and non-zero parallel ${\rm spin}^c$ spinors are stable (X. Dai, X. Wang, and G. Wei, 2007).

- Einstein manifolds with negative sectional curvature (N. Koiso, 1978).
- Many compact irreducible symmetric spaces, e.g. S^n , $\mathbb{C}P^n$, $\mathbb{H}P^n$, and all compact simple Lie groups except $\mathrm{Sp}(n)$, $n \geq 2$ and $\mathrm{SU}(n)$, $n \geq 3$ (N. Koiso, 1980).
- Compact manifolds with non-zero parallel spinors (X. Dai, X. Wang, and G. Wei, 2005). Therefore, so far all known examples of compact Ricci flat manifold are stable.
- Compact Einstein manifolds with non-positive scalar curvature and non-zero parallel spin^c spinors are stable (X. Dai, X. Wang, and G. Wei, 2007). In particular, compact Kähler-Einstein manifolds with non-positive scalar curvature are stable.

- Einstein manifolds with negative sectional curvature (N. Koiso, 1978).
- Many compact irreducible symmetric spaces, e.g. S^n , $\mathbb{C}P^n$, $\mathbb{H}P^n$, and all compact simple Lie groups except $\mathrm{Sp}(n)$, $n \geq 2$ and $\mathrm{SU}(n)$, $n \geq 3$ (N. Koiso, 1980).
- Compact manifolds with non-zero parallel spinors (X. Dai, X. Wang, and G. Wei, 2005). Therefore, so far all known examples of compact Ricci flat manifold are stable.
- Compact Einstein manifolds with non-positive scalar curvature and non-zero parallel spin^c spinors are stable (X. Dai, X. Wang, and G. Wei, 2007). In particular, compact Kähler-Einstein manifolds with non-positive scalar curvature are stable.
- Complete Riemannian manifolds with imaginary Killing spinors (K. Kröncke, 2017, W., 2017).

Examples of \widetilde{S} -linearly unstable Einstein metrics

• $Sp(n), Sp(n)/U(n), n \ge 2$ (N. Koiso, 1980).

Examples of $\widetilde{\mathbf{S}}$ -linearly unstable Einstein metrics

- $\mathrm{Sp}(n), \mathrm{Sp}(n)/\mathrm{U}(n), n \geq 2$ (N. Koiso, 1980).
- ullet Jensen spheres metric on S^{4q+3} (G. Jensen, 1973).

- $\mathrm{Sp}(n), \mathrm{Sp}(n)/\mathrm{U}(n), n \geq 2$ (N. Koiso, 1980).
- Jensen spheres metric on S^{4q+3} (G. Jensen, 1973).
- Any compact Kähler-Einstein manifold M of positive scalar curvature with dim $H^{1,1}(M) \ge 2$ (i.e. $b_2 \ge 2$) is unstable (H. Cao, R. Hamilton, and T. Ilmanen, 2004).

- $\mathrm{Sp}(n), \mathrm{Sp}(n)/\mathrm{U}(n), n \geq 2$ (N. Koiso, 1980).
- Jensen spheres metric on S^{4q+3} (G. Jensen, 1973).
- Any compact Kähler-Einstein manifold M of positive scalar curvature with dim $H^{1,1}(M) \ge 2$ (i.e. $b_2 \ge 2$) is unstable (H. Cao, R. Hamilton, and T. Ilmanen, 2004).
- Warped product Einstein manifolds with dimension $n \le 6$ are unstable. (W. Batat, S. Hall, T. Murphy, 2017).

Some other variational characterizations of Einstein metrics

Einstein metrics with positive Ricci curvature are certain critical points of Perelman's ν -entropy.

Some other variational characterizations of Einstein metrics

Einstein metrics with positive Ricci curvature are certain critical points of Perelman's ν -entropy.

The second variation formula of ν -entropy at Einstein metrics for $h \in \ker \operatorname{tr}_g \cap \ker \delta_g$ is the same as the second variation formula of the normalized total scalar curvature (up to a positive constant factor).

Some other variational characterizations of Einstein metrics

Einstein metrics with positive Ricci curvature are certain critical points of Perelman's *v*-entropy.

The second variation formula of ν -entropy at Einstein metrics for $h \in \ker \operatorname{tr}_g \cap \ker \delta_g$ is the same as the second variation formula of the normalized total scalar curvature (up to a positive constant factor).

However, unlike the case of the $\tilde{\mathbf{S}}$ -functional, the second variation is no longer always positive on conformal change directions. Actually, Einstein metrics could be local maximum points of the ν -entropy.

ν -(linear) stability

Definition: A closed Einstein manifold (M,g) with Einstein constant $\Lambda > 0$ is (1) ν -stable if g is a local maximizer of the ν -entropy;

ν -(linear) stability

Definition: A closed Einstein manifold (M,g) with Einstein constant $\Lambda > 0$ is

- (1) ν -stable if g is a local maximizer of the ν -entropy;
- (2) ν -linearly stable if the second variation of the ν -entropy is negative semi-definite on $C^{\infty}(M)g \oplus (\ker \operatorname{tr}_g \cap \ker \delta_g)$.

The corresponding notions of instability are given by negation.

ν -(linear) stability

Definition: A closed Einstein manifold (M,g) with Einstein constant $\Lambda > 0$ is

- (1) ν -stable if g is a local maximizer of the ν -entropy;
- (2) ν -linearly stable if the second variation of the ν -entropy is negative semi-definite on $C^{\infty}(M)g \oplus (\ker \operatorname{tr}_g \cap \ker \delta_g)$.

The corresponding notions of instability are given by negation.

By H. Cao-C. He's work,

$$\nu\text{-linear stability} \Longleftrightarrow \begin{cases} \langle \nabla^*\nabla h - 2\mathring{R}h, h \rangle_{L^2} \geq 0, & \forall \ h \in \ker \operatorname{tr}_g \cap \ker \delta_g, \\ \lambda_1(M,g) \geq 2\Lambda, \end{cases}$$

where $\lambda_1(M,g)$ is the first non-zero eigenvalue of the Laplace operator on (M,g).

Relation with dynamic stability

By the work of K. Kröncke (2015), ν -stability is equivalent to dynamical stability along normalized Ricci flow.

Relation with dynamic stability

By the work of K. Kröncke (2015), ν -stability is equivalent to dynamical stability along normalized Ricci flow.

A compact Einstein manifold (M,g) is called dynamically unstable if there exists a non-trivial normalized Ricci flow defined on $(-\infty,0]$ which converges modulo diffeomorphism to g as $t\to -\infty$.

Relation with dynamic stability

By the work of K. Kröncke (2015), ν -stability is equivalent to dynamical stability along normalized Ricci flow.

A compact Einstein manifold (M,g) is called dynamically unstable if there exists a non-trivial normalized Ricci flow defined on $(-\infty,0]$ which converges modulo diffeomorphism to g as $t\to -\infty$.

 $\widetilde{\mathbf{S}}\text{-linearly unstable} \Longrightarrow \nu\text{-linearly unstable} \Longrightarrow \nu\text{-unstable} \Longleftrightarrow$ dynamically unstable.

A spinor σ is called a Killing spinor with the Killing constant $\mu \neq 0$, if

$$\nabla_X^S \sigma = \mu X \cdot \sigma, \tag{9}$$

for all vector fields X on M.

A spinor σ is called a Killing spinor with the Killing constant $\mu \neq 0$, if

$$\nabla_X^S \sigma = \mu X \cdot \sigma, \tag{9}$$

for all vector fields X on M.

If the constant $\mu = 0$ in (9), then σ is called a parallel spinor.

A spinor σ is called a Killing spinor with the Killing constant $\mu \neq 0$, if

$$\nabla_X^S \sigma = \mu X \cdot \sigma, \tag{9}$$

for all vector fields X on M.

If the constant $\mu = 0$ in (9), then σ is called a parallel spinor.

If a manifold (M^n, g) admits a Killing spinor σ with Killing constant μ , then

$$Ric_g = 4\mu^2(n-1)g, \tag{10}$$

i.e. g is an Einstein metric with scalar curvature $4n(n-1)\mu^2$. This implies that μ can only be real or purely imaginary, since scalar curvature is real.

A spinor σ is called a Killing spinor with the Killing constant $\mu \neq 0$, if

$$\nabla_X^S \sigma = \mu X \cdot \sigma, \tag{9}$$

for all vector fields X on M.

If the constant $\mu = 0$ in (9), then σ is called a parallel spinor.

If a manifold (M^n,g) admits a Killing spinor σ with Killing constant μ , then

$$Ric_g = 4\mu^2(n-1)g, \tag{10}$$

i.e. g is an Einstein metric with scalar curvature $4n(n-1)\mu^2$. This implies that μ can only be real or purely imaginary, since scalar curvature is real.

A non-zero Killing spinor is real (resp. imaginary) if its Killing constant is real (resp. purely imaginary).

A spinor σ is called a Killing spinor with the Killing constant $\mu \neq 0$, if

$$\nabla_X^S \sigma = \mu X \cdot \sigma, \tag{9}$$

for all vector fields X on M.

If the constant $\mu = 0$ in (9), then σ is called a parallel spinor.

If a manifold (M^n, g) admits a Killing spinor σ with Killing constant μ , then

$$Ric_g = 4\mu^2(n-1)g, \tag{10}$$

i.e. g is an Einstein metric with scalar curvature $4n(n-1)\mu^2$. This implies that μ can only be real or purely imaginary, since scalar curvature is real.

A non-zero Killing spinor is real (resp. imaginary) if its Killing constant is real (resp. purely imaginary).

Professor Thomas Friedrich initiated the mathematical investigation of Killing spinors in 1980.



Manifolds with real Killing spinors

By the works of T. Friedrich (1981), O. Hijazi (1986), T. Friedrich-I. Kath (1989), R. Grunewald (1990), C. Bär (1993), a complete simply-connected Riemannian manifold of dimension n with a Real Killing spinor and Einstein constant n-1 is isometric one of the following:

- (1) round sphere S^n , if n is even and $n \neq 6$;
- (2) strictly nearly Kähler manifold, if n = 6;
- (3) Sasaki-Einstein manifold, if n is odd and $n \neq 7$;
- (4) nearly parallel G_2 manifold (including Sasaki-Einstein and 3-Sasaki), if n=7.

Homogeneous nearly Kähler manifolds

Proposition (W. - Wang, 2018)

The only ν -stable simply connected homogeneous strictly nearly Kähler 6-manifold is S^6 with the round metric.

Homogeneous nearly Kähler manifolds

Proposition (W. – Wang, 2018)

The only ν -stable simply connected homogeneous strictly nearly Kähler 6-manifold is S^6 with the round metric.

By the work of J.-B. Butruille (2005), a simply connected homogeneous strictly nearly Kähler 6-manifold is one of the following:

- (1) $S^6 = G_2/SU(3)$,
- (2) $(SU(2) \times SU(2) \times SU(2))/\Delta SU(2)$,
- (3) $\mathbb{C}P^3 = \text{Sp}(2)/(\text{Sp}(1) \times \text{U}(1)),$
- (4) $SU(3)/T^2$,

each equipped with a unique invariant nearly Kähler structure.

Instability of nearly Kähler 6-manifoolds

Theorem (Semmelmann – W. – Wang, 2019)

A complete strict nearly Kähler 6-manifold with either 2^{nd} or 3^{rd} Betti number nonzero is $\widetilde{\mathbf{S}}$ -linearly unstable, and therefore ν -linearly unstable.

Instability of nearly Kähler 6-manifoolds

Theorem (Semmelmann – W. – Wang, 2019)

A complete strict nearly Kähler 6-manifold with either 2^{nd} or 3^{rd} Betti number nonzero is $\widetilde{\mathbf{S}}$ -linearly unstable, and therefore ν -linearly unstable.

Main ingredients in the proof:

- M. Verbitsky's Hodge decomposition of Harmonic forms on nearly Kähler 6-manifolds (also proved by L. Foscolo in a different way).
- A. Moroianu and U. Semmelmann's work about the standard Lapace operator on nearly Kähler manifolds.

Instability of nearly Kähler 6-manifoolds

Theorem (Semmelmann – W. – Wang, 2019)

A complete strict nearly Kähler 6-manifold with either 2^{nd} or 3^{rd} Betti number nonzero is $\tilde{\mathbf{S}}$ -linearly unstable, and therefore ν -linearly unstable.

Main ingredients in the proof:

- M. Verbitsky's Hodge decomposition of Harmonic forms on nearly Kähler 6-manifolds (also proved by L. Foscolo in a different way).
- A. Moroianu and U. Semmelmann's work about the standard Lapace operator on nearly Kähler manifolds.

Corollary

If a complete simply connected strict nearly Kähler manifolds is $\widetilde{\mathbf{S}}$ -linearly stable, then it is a rational homology sphere. In particular, if $H_2(M,\mathbb{Z})$ has no torsion, then it is diffeomorphic to S^6 .

Nearly parallel G₂ manifolds

Classification by Friedrich-Kath-Moroianu-Semmelmann (1997):

A compact, simply connected, homogeneous nearly parallel G_2 manifold with only 1 linearly independent Killing spinor is isometric to one of the followings

- (1) S^7 with the Jensen's metric ($\widetilde{\mathbf{S}}$ -linearly unstable);
- (2) Aloff-Wallach spaces $N_{k,l} = \mathrm{SU}(3)/U_{k,l}$, where k,l are relatively prime integers with $(k,l) \neq (1,1)$ and $U_{k,l}$ is the circle $\mathrm{diag}(e^{2\pi i k \theta}, e^{2\pi i l \theta}, e^{-2\pi i (k+l) \theta}) \in \mathrm{SU}(3)$, with invariant Einstein metrics;
- (3) the isotropy irreducible space $\mathrm{Sp}(2)/\mathrm{SU}(2)$, where the embedding of $\mathrm{SU}(2)$ is via the irreducible 4-dimensional symplectic representation.

Instability of homogeneous Einstein metrics on Aloff-Wallach spaces $N_{k,l}$

By the works of M. Wang (1982), Castellani-Romans (1984), Page-Pope (1984), Kawalski-Vlasek (1993), and Nikonorov (2004), each of $N_{k,l}$ admit two SU(3)-invariant Einstein metrics, up to isometry.

Instability of homogeneous Einstein metrics on Aloff-Wallach spaces $N_{k,l}$

By the works of M. Wang (1982), Castellani-Romans (1984), Page-Pope (1984), Kawalski-Vlasek (1993), and Nikonorov (2004), each of $N_{k,l}$ admit two $\mathrm{SU}(3)$ -invariant Einstein metrics, up to isometry.

Theorem (W. – Wang, 2018)

The invariant Einstein metrics on $N_{k,l}$ are all **S**-linearly unstable, and therefore ν -linearly unstable.

Definition 1 of Sasaki Manifolds: (M^n,g) is a Sasaki manifold if the cone $(\mathbb{R}_+ \times M^n, dr^2 + r^2g)$ is Kähler, where $\mathbb{R}_+ = (0, +\infty)$, and r is coordinate on \mathbb{R}_+ .

Definition 1 of Sasaki Manifolds: (M^n, g) is a Sasaki manifold if the cone $(\mathbb{R}_+ \times M^n, dr^2 + r^2g)$ is Kähler, where $\mathbb{R}_+ = (0, +\infty)$, and r is coordinate on \mathbb{R}_+ .

Definition 2 of Sasaki Manifolds: (M^n,g) is a Sasaki manifold if there exists a Killing vector filed ξ of unit length on M^n so that the Riemann curvature satisfies the condition

$$R_{X\xi}Y = -g(\xi, Y)X + g(X, Y)\xi, \tag{11}$$

for any pair of vector fields X and Y on M^n .

Definition 1 of Sasaki Manifolds: (M^n, g) is a Sasaki manifold if the cone $(\mathbb{R}_+ \times M^n, dr^2 + r^2g)$ is Kähler, where $\mathbb{R}_+ = (0, +\infty)$, and r is coordinate on \mathbb{R}_+ .

Definition 2 of Sasaki Manifolds: (M^n,g) is a Sasaki manifold if there exists a Killing vector filed ξ of unit length on M^n so that the Riemann curvature satisfies the condition

$$R_{X\xi}Y = -g(\xi, Y)X + g(X, Y)\xi, \tag{11}$$

for any pair of vector fields X and Y on M^n . Moreover, (M^n, g) is a regular Sasaki manifold if ξ is a regular vector field.

Definition 1 of Sasaki Manifolds: (M^n, g) is a Sasaki manifold if the cone $(\mathbb{R}_+ \times M^n, dr^2 + r^2g)$ is Kähler, where $\mathbb{R}_+ = (0, +\infty)$, and r is coordinate on \mathbb{R}_+ .

Definition 2 of Sasaki Manifolds: (M^n,g) is a Sasaki manifold if there exists a Killing vector filed ξ of unit length on M^n so that the Riemann curvature satisfies the condition

$$R_{X\xi}Y = -g(\xi, Y)X + g(X, Y)\xi, \tag{11}$$

for any pair of vector fields X and Y on M^n . Moreover, (M^n, g) is a regular Sasaki manifold if ξ is a regular vector field.

From (11), one can easily see that on a Sasaki-Einstein manifold (M^n, g) of dimension n

$$Ric_g = (n-1)g. (12)$$



Regular Sasaki-Einstein manifolds

Let (B^{2p}, G, J) be a Kähler-Einstein manifold of real dimension 2p, with the Kähler form $\Omega = G(\cdot, J \cdot)$, and $Ric_G = (2p + 2)G$.

Then let $\pi:M^{2p+1}\to B^{2p}$ be a principal S^1 -bundle with a connection η with the curvature form $d\eta=2\pi^*\Omega$.

Let ξ be a vertical vector field on M^{2p+1} , generated by S^1 -action, such that $\eta(\xi)=1$.

We define a Riemannian metric on M^{2p+1} as

$$g(X,Y) = G(\pi_*X, \pi_*Y) + \eta(X)\eta(Y),$$
 (13)

for vector field X and Y on M^{2p+1} .

Then (M^{2p+1},g,ξ) is a regular Sasaki-Einstein manifold.

Einstein operator on regular Sasaki-Einstein manifolds

Proposition (W., 2016)

$$\int_{M} \langle (\nabla^{g})^{*} \nabla^{g} \tilde{h} - 2\mathring{R}^{g} \tilde{h}, \tilde{h} \rangle d\operatorname{vol}_{g}$$

$$= \int_{B} (\langle (\nabla^{G})^{*} \nabla^{G} h - 2\mathring{R}^{G} h, h \rangle + 4\langle h, h \rangle + 4\langle h \circ J, h \rangle) d\operatorname{vol}_{G},$$
for any $h \in C^{\infty}(B, S^{2}(B))$, where $\tilde{h} = \pi^{*} h \in C^{\infty}(M, S^{2}(M))$.

Einstein operator on regular Sasaki-Einstein manifolds

Proposition (W., 2016)

$$\int_{M} \langle (\nabla^{g})^{*} \nabla^{g} \tilde{h} - 2\mathring{R}^{g} \tilde{h}, \tilde{h} \rangle d \operatorname{vol}_{g}$$

$$= \int_{B} (\langle (\nabla^{G})^{*} \nabla^{G} h - 2\mathring{R}^{G} h, h \rangle + 4 \langle h, h \rangle + 4 \langle h \circ J, h \rangle) d \operatorname{vol}_{G}, \tag{14}$$

for any $h \in C^{\infty}(B, S^2(B))$, where $\tilde{h} = \pi^* h \in C^{\infty}(M, S^2(M))$.

Corollary

The regular Sasaki-Einstein manifold (M^{2p+1},g) is unstable, if one of the following conditions holds:

(1) there exists a traceless transverse symmetric 2-tensor h on base manifold B^{2p} such that $\int_{\mathcal{B}} \langle (\nabla^G)^* \nabla^G h - 2\mathring{R}^G h, h \rangle d\mathrm{vol}_G < -8 \int_{\mathcal{B}} \langle h, h \rangle d\mathrm{vol}_G;$

Einstein operator on regular Sasaki-Einstein manifolds

Proposition (W., 2016)

$$\int_{M} \langle (\nabla^{g})^{*} \nabla^{g} \tilde{h} - 2\mathring{R}^{g} \tilde{h}, \tilde{h} \rangle d \operatorname{vol}_{g}$$

$$= \int_{B} (\langle (\nabla^{G})^{*} \nabla^{G} h - 2\mathring{R}^{G} h, h \rangle + 4 \langle h, h \rangle + 4 \langle h \circ J, h \rangle) d \operatorname{vol}_{G}, \tag{14}$$

for any $h \in C^{\infty}(B, S^2(B))$, where $\tilde{h} = \pi^* h \in C^{\infty}(M, S^2(M))$.

Corollary

The regular Sasaki-Einstein manifold (M^{2p+1}, g) is unstable, if one of the following conditions holds:

- (1) there exists a traceless transverse symmetric 2-tensor h on base manifold B^{2p} such that $\int_{\mathcal{B}} \langle (\nabla^G)^* \nabla^G h 2\mathring{R}^G h, h \rangle d\mathrm{vol}_G < -8 \int_{\mathcal{B}} \langle h, h \rangle d\mathrm{vol}_G;$
- (2) the base Kähler-Einstein manifold (B^{2p},G,J) has $dim H^{1,1}(B) \geq 2$.



Instability of regular Sasaki-Einstein manifolds with Hermitian symmetric space bases

Theorem (W. – Wang, 2018)

The following simply connected regular Sasaki Einstein manifold are ν -linearly unstable from conformal variations:

- (1) SO(p+2)/SO(p), $p \ge 3$, circle bundle over the complex quadric $SO(p+2)/(SO(p) \times SO(2))$;
- (2) E₆/Spin(10), and E₇/E₆, which are respectively circle bundles over the hermitian symmetric spaces E₆/(Spin(10 · U(1)) and E₇/(E₆ · U(1));
- (3) $SU(p+2)/(SU(p) \times SU(2)), p \ge 2$, a circle bundle over the complex Grassmannian $SU(p+2)/S(U(p) \times U(2))$.

Moreover, the Stiefel manifolds in (a) above are also $\widetilde{\mathbf{S}}$ -linearly unstable, and for $k \geq 4$, $\mathrm{Sp}(k)/\mathrm{SU}(k)$, which are circle bundles over $\mathrm{Sp}(k)/\mathrm{U}(k)$, are $\widetilde{\mathbf{S}}$ -linearly unstable, and so ν -linearly unstable.

THANK YOU!